

Three Topics on Normal Modes: Barotropic Equatorial Trapping and the Effective Lamb's Parameter, Kelvin Solitons and Corner Waves and Hough Eigenvalue Point Clouds

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FIRST THEME: Equatorial Trapping is
Controlled by BOTH Lamb's Parameter and
Zonal Wavenumber

$$\epsilon \equiv \frac{4\Omega a^2}{gH} \text{ [LAMB'S PARAMETER]}$$

$$\Omega = 2\pi / 84,600 \text{ s}, \quad a = \text{earth's radius}$$
$$g = 9.8 \text{ m/s} \quad H = \text{equivalent depth}$$

$s \equiv$ LONGITUDINAL WAVENUMBER,
an INTEGER

Equatorial Beta-Plane: Asymptotic Approximation by Hermite Functions

- Orthodoxy: $v_n \sim \psi_n(\epsilon^{1/4}\mu)$,
 $\psi_n(\gamma) = \exp(-[1/2]\gamma^2)H_n(\gamma)$
- Boyd (*J. Atmos. Sci.*, 1985) argued that for Rossby waves,

$$\boxed{\epsilon \rightarrow \epsilon + s^2}$$

- Boyd & Zhou (*J. Atmos. Sci.*, 2008) extended to Kelvin waves
- Argument applies to SPHERICAL HARMONICS & PROLATE SPHEROIDAL FUNCTIONS as well as HOUGH FUNCTIONS

Prolate Spheroidal Illustration

$$\frac{d^2\psi}{d\phi^2} - \frac{d\psi}{d\phi} + \left\{ \chi_n - \frac{s^2}{\cos^2(\phi)} - \epsilon \sin(\phi)^2 \right\} \psi = 0$$

$$\Downarrow \cos(\phi) \rightarrow 1, \sin(\phi) \rightarrow \phi \Downarrow$$

$$\frac{d^2v}{d\phi^2} + \left\{ \chi_n - s^2 - \epsilon \phi^2 \right\} v = 0$$

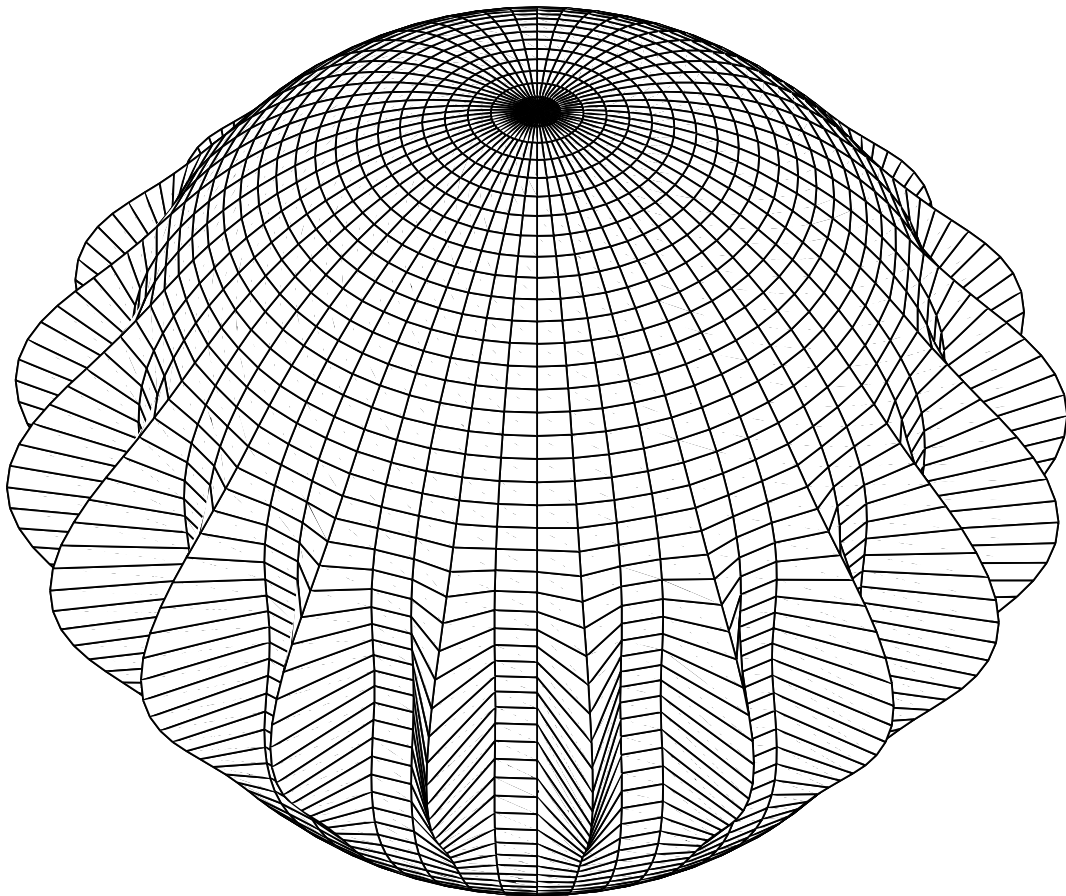
$$\begin{aligned} 1 / \cos(\phi)^2 &= 1 / \left\{ 1 - \sin^2(\phi) \right\} \\ &= 1 + \sin^2(\phi) + O(\sin^4(\phi)) \\ &= 1 + \phi^2 + O(\phi^4) \end{aligned}$$

$$\begin{aligned} &-\frac{s^2}{\cos^2(\phi)} - \epsilon \sin(\phi)^2 \quad (1) \\ \Rightarrow &-s^2 - (\epsilon + s^2) \sin(\phi)^2 \end{aligned}$$

Barotropic ($\epsilon = 0$) Kelvin Waves

High zonal wavenumber Kelvin are equatorial modes even for $\epsilon = 0$

Barotropic Kelvin, $\varepsilon=0$, $s=20$

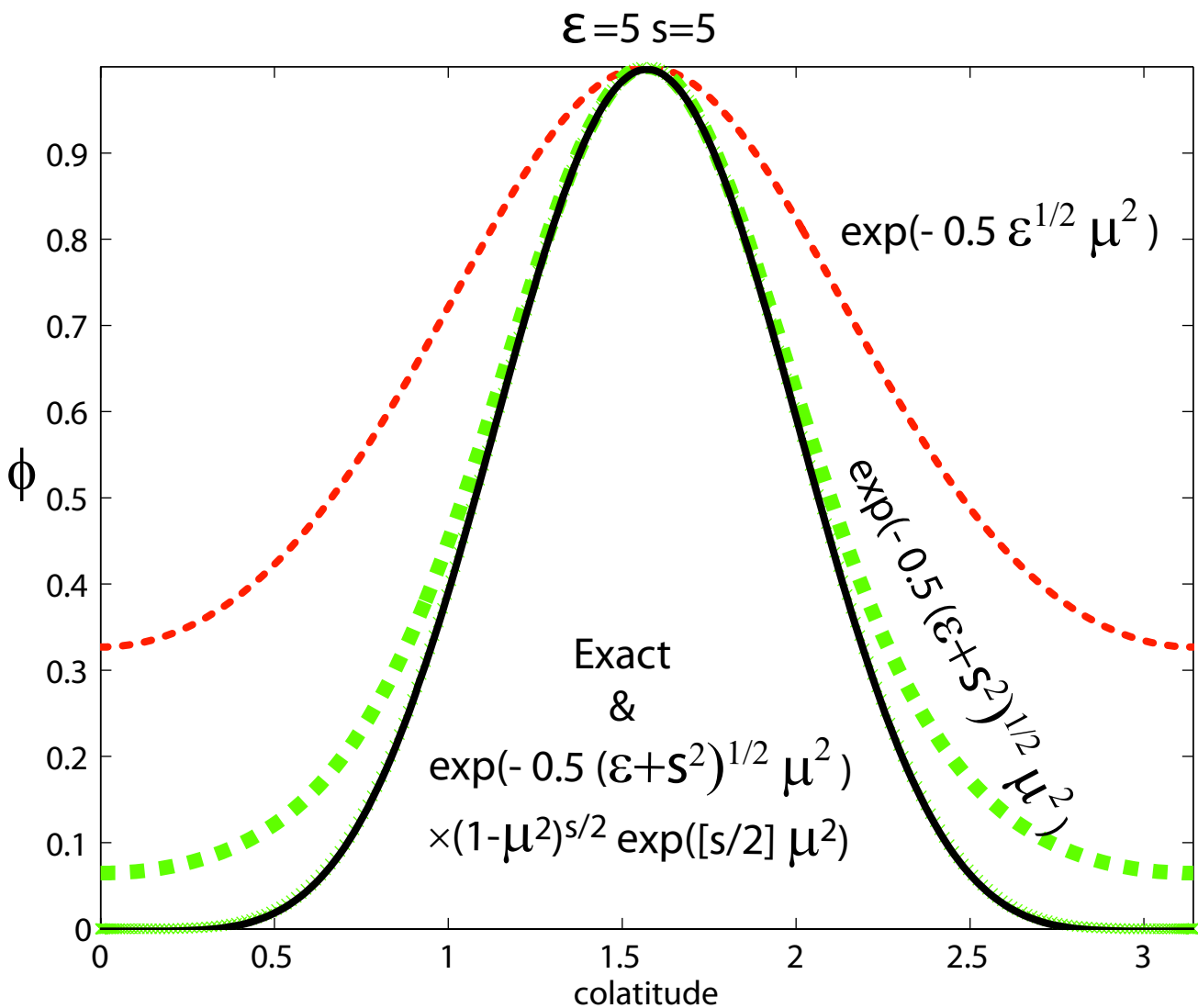


Boyd-Zhou Kelvin approximation is

$$\phi \approx (1-\mu^2)^{s/2} \exp((s/2)\mu^2) \times \exp(- (1/2)\sqrt{\epsilon + s^2}\mu^2)$$

($\mu = \sin(\text{latitude})$)

Kelvin & approx. are solid black [graphically indistinguishable]



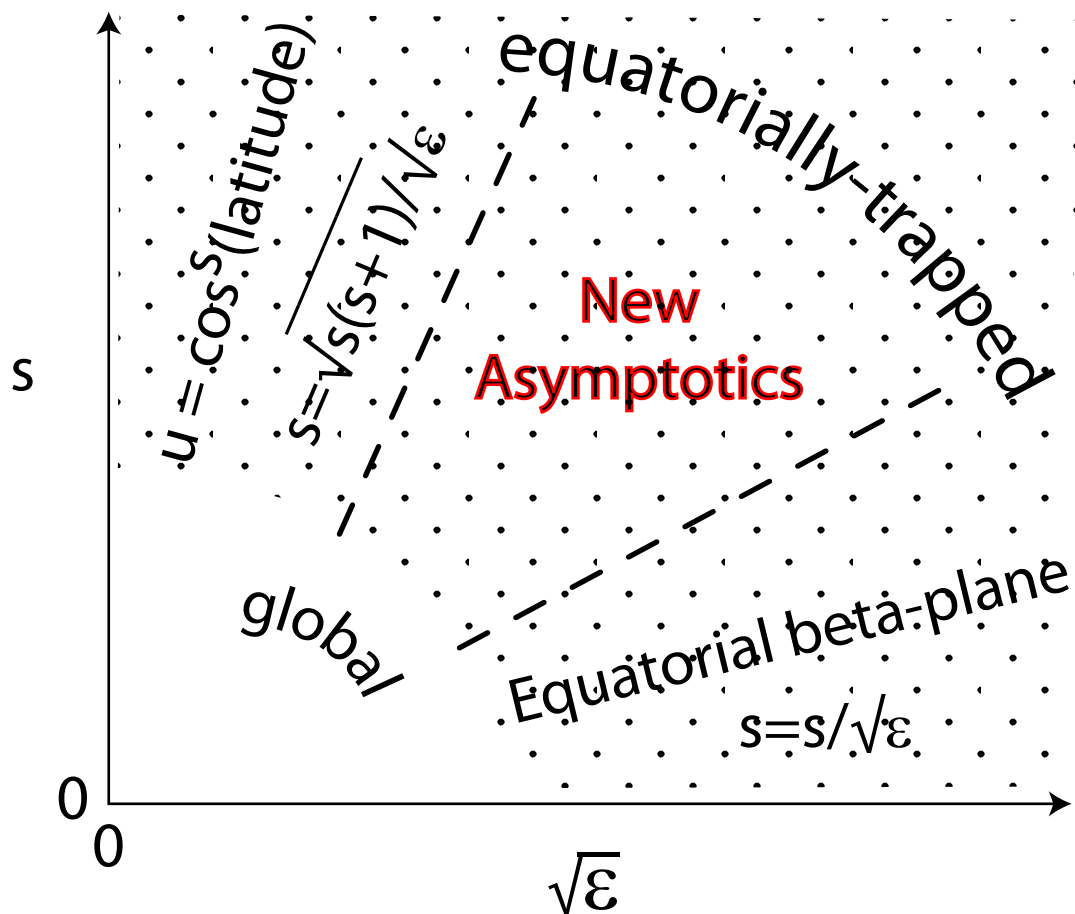
Uniform Validity

- New approximation is uniformly valid for

$$\sqrt{s^2 + \epsilon} \gg 1$$

(shaded in figure)

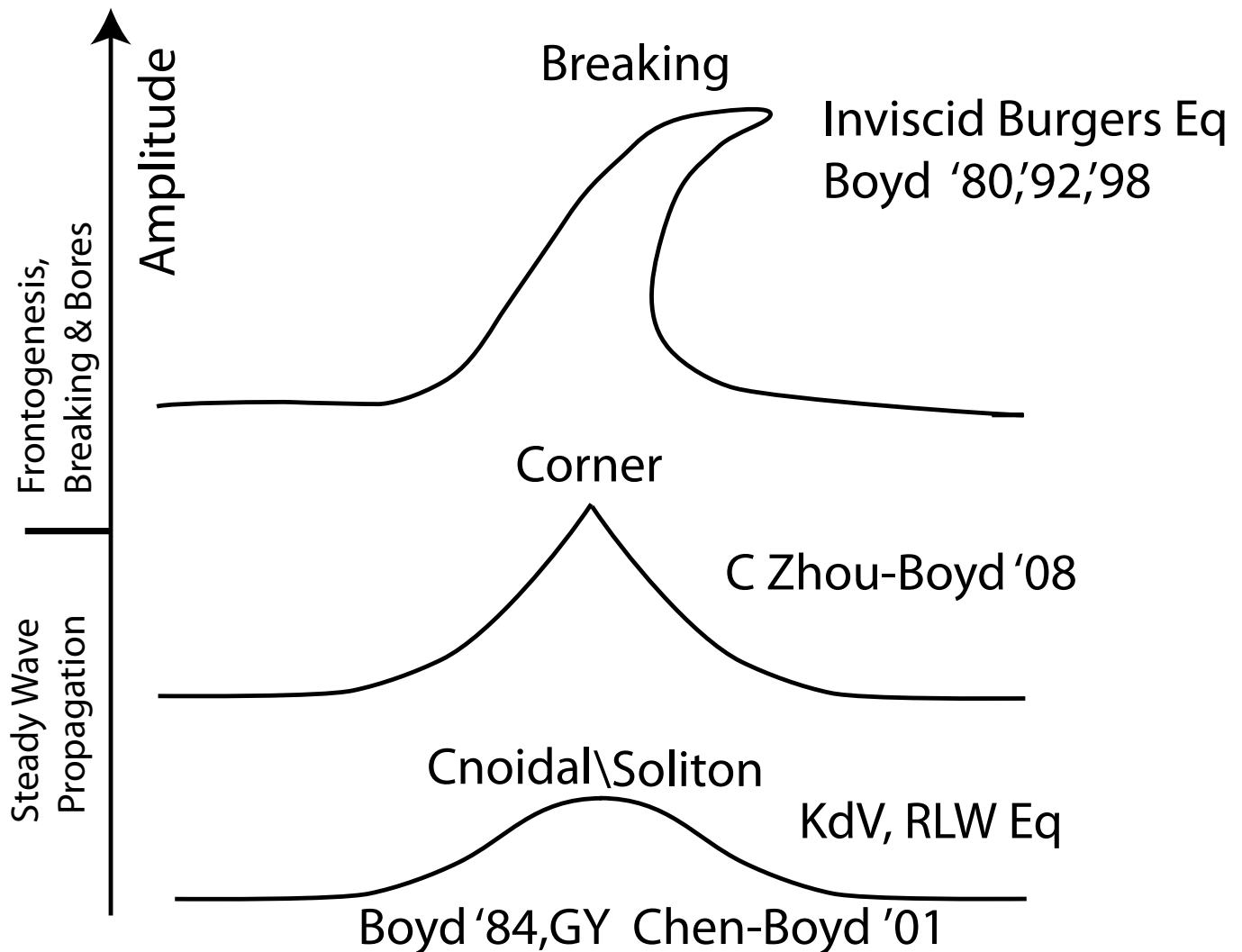
- Though not strictly valid when both s and ϵ are $O(1)$, it is not a bad approximations



SECOND THEME: NONLINEAR KELVIN DYNAMICS

KELVIN MODE

CCB Scenario: Cnoidal/Corner/Breaking



Definition 1 (Corner Wave) *A corner wave is a steadily traveling nonlinear wave in which the wave height function $u(x - ct)$ has a maximum which is a slope discontinuity.*

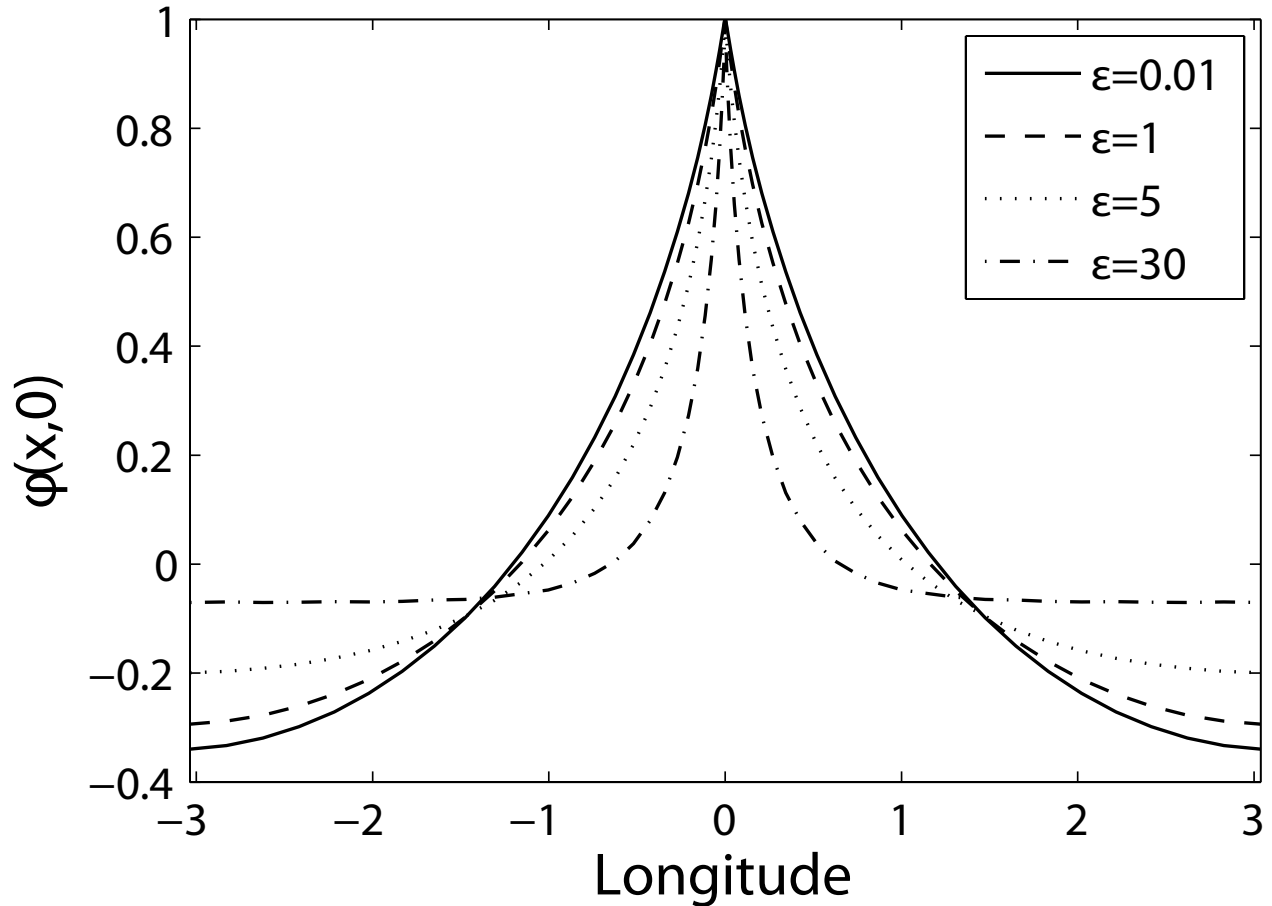
Table 1: Examples of Systems with Corner Waves and the CCB Scenario

Eq. or Wave Name	Equations
Non-equatorial	
Equatorial Waves Barotropic Mode	$K_t + KK_x = \gamma \{ Y(t)e^{ix} + \bar{Y}(t)e^{-ix} \};$ $Y_t = -\gamma \hat{K}(x = 1, t)$
Equatorial Waves Baroclinic Mode	3 coupled PDEs in (x, t)
Resonant Triads, One Nondispersive	$u_t + uu_x = 2\text{Re}(ikab \exp(-ikx));$ $a_t = -i\omega_a \bar{b} \hat{u}_k; b_t = -i\omega_b a \hat{u}_k$
Equatorial Kelvin (4-mode Model)	4 coupled PDEs in x, t
Equatorial Kelvin (Shallow Water)	3 coupled PDEs in x, y, t (Shallow Water Eqs.)
Non-equatorial	
Surface Irrotational Water Waves	Euler equations in x, z
Boundary Waves on Vortex Patches	Two-space-dimensional Euler equations (x, y)
Camassa-Holm	$u_t - u_{xxt} + (2\kappa + 3u - 2u_{xx})u_x - uu_{xxx} = 0$
Ostrovsky-Hunter	$(u_t + uu_x)_x = u$
Gabov/ Shefter-Rosales	$(u_t + uu_x)_x = \int_0^{2\pi} \cos(x - y) u(y) dy$
Whitham	$(u_t + uu_x)_x = pb^2 \times$ $\left\{ u - \int_0^{2\pi} \frac{b \cosh(b \{ X-y - \pi \})}{2 \sinh(b\pi)} u(y) dy \right\}$

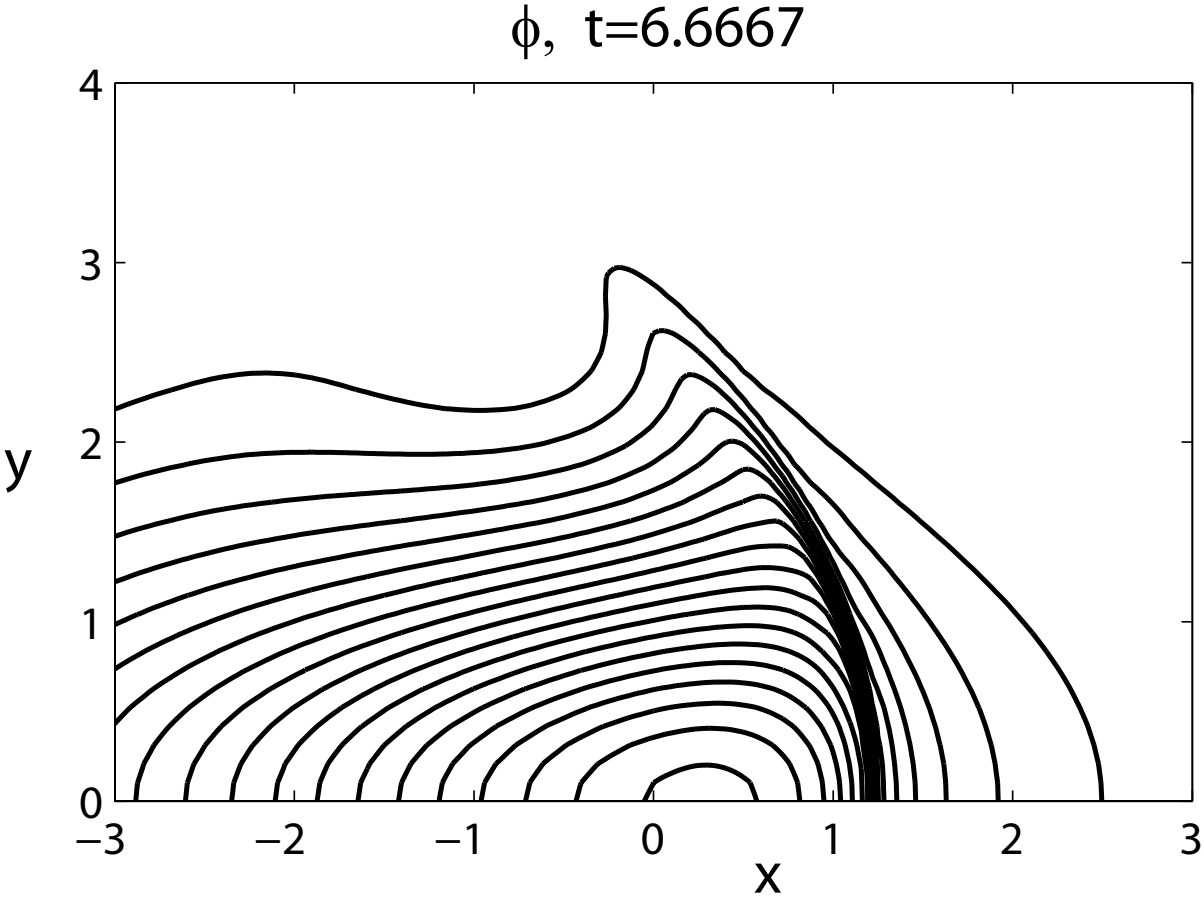
Corner waves for different values of Lamb's parameter ϵ

CORNER WAVE is a POINT SINGULARITY
NOT a CREASE
NOT a CONE

$$\varphi(x, \text{equator}) / \varphi_{00} \quad s=1$$



Kelvin front CURVES because of resonance with gravity waves



HOUGH POINT CLOUDS INTO POLYNOMIALS

Hough's spherical harmonic Galerkin algorithm, with Longuet-Higgins's improvements, is very fast and spectrally accurate.

Mode classification is NOT a SLAM DUNK

Galerkin method generates POINT CLOUD: eigenvalues at discrete ϵ .

Desired: CONTINUOUS BRANCHES

Other complications:

Kelvin mode \Rightarrow GRAVITY WAVE as $\epsilon \rightarrow 0$

Yanai mode is "MIXED ROSSBY-GRAVITY"

Number of interior zeros may change with ϵ



Making Friends with Special Functions

CONCEPTUAL, QUALITATIVE:

Never-Out-of-Date Paradigms: Theorems,
Asymptotics & Graphs

NUMERICAL:

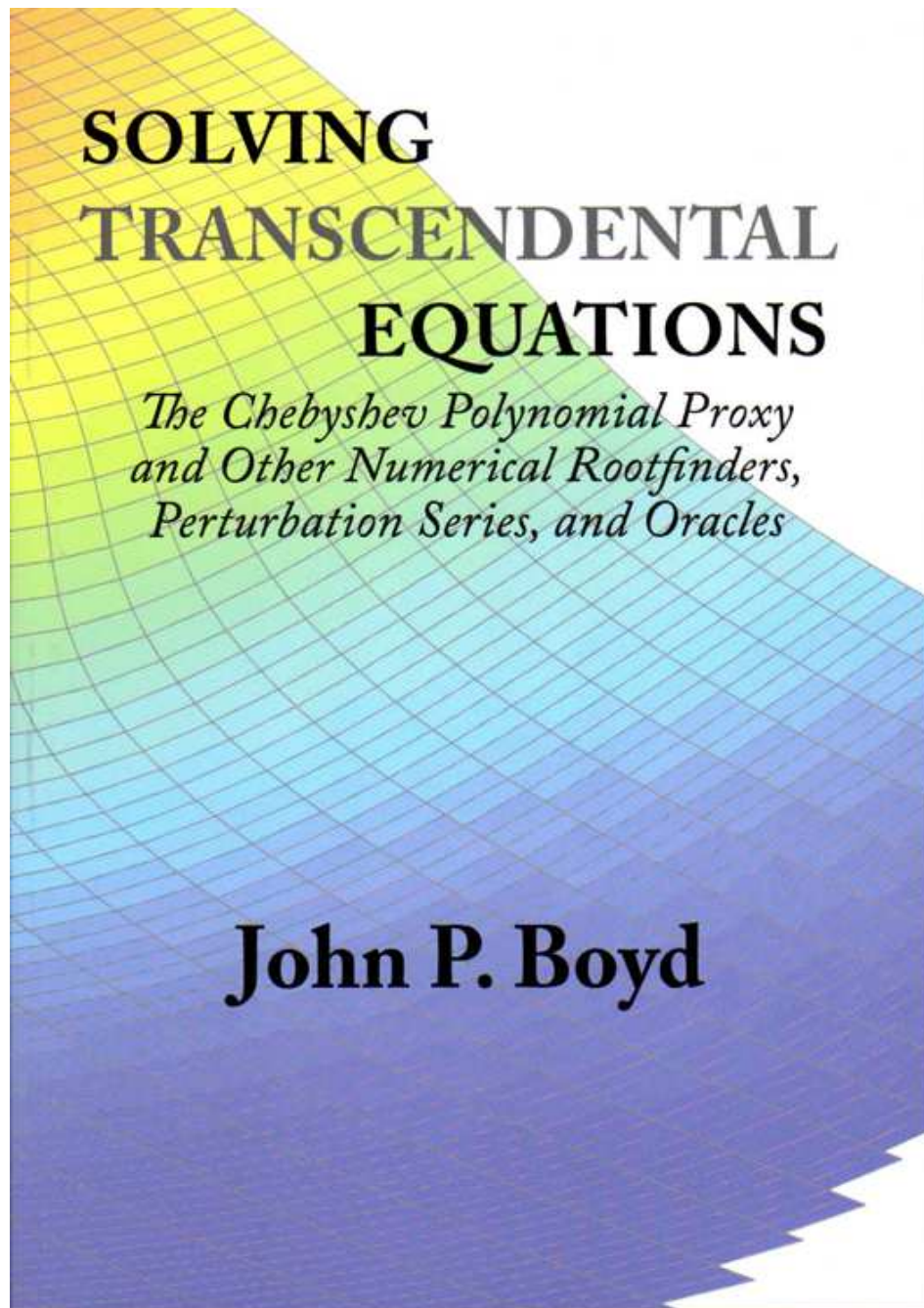
Ancient Paradigm: Tables

Newer Paradigm: Perturbation Series & Chebyshev Series

Emerging Paradigm: Matlab Code

Spherical harmonic Galerkin discretizations are tridiagonal (ϵ is eigenvalue) or otherwise very sparse. Power method allows very fast computation of a chosen mode for arbitrary parameter values without the need to compute all other modes if a Never-Failing-Initialization available.

Alternatives to Never-Failing-Initialization
Continuation, Davidenko Equation, etc.,
WORK but MANY POTENTIAL PROBLEMS
Discussed in many references including:



Never-Failing-Initialization: Seven Series Options

small ε

spherical
harmonics

large ε

Hermite
functions

Pade from
small ε

Pade from
large ε

all ε

TWO-POINT
PADE

Rational Chebyshev (TL series)
for each s

Double Rational Chebyshev

(\mathcal{E}, s)

Two-Point Padé Approximants

Example: Kelvin wave

The existence of such approximations suggests a unity of structure and identity in the Kelvin mode over all of $\epsilon \in [0, \infty]$.

Linear Polynomial/Linear Polynomial in $\sqrt{\epsilon}$ Matches
 (i) $\epsilon = 0$ limit (ii) two terms in $1/\sqrt{\epsilon}$:

$$c_{[1/1]}^{two-point} = \left(\sqrt{\frac{s+1}{s}} + 4\epsilon^{1/2} \sqrt{\frac{s+1}{s}} - 4\epsilon^{1/2} \right) \left(1 + 4\epsilon^{1/2} \sqrt{\frac{s+1}{s}} - 4\epsilon^{1/2} \right)^{-1} \quad (2)$$

The next Kelvin approximation $c_{[2/2]}^{two-point}$ matches the first three terms of the large- ϵ expansion and two terms of the small- ϵ series [not shown]

The maximum relative error of the two-point Padé $c_{[2/2]}^{two-point}$ for Kelvin mode is only 0.0184 over all of $\epsilon \in [0, \infty]$.

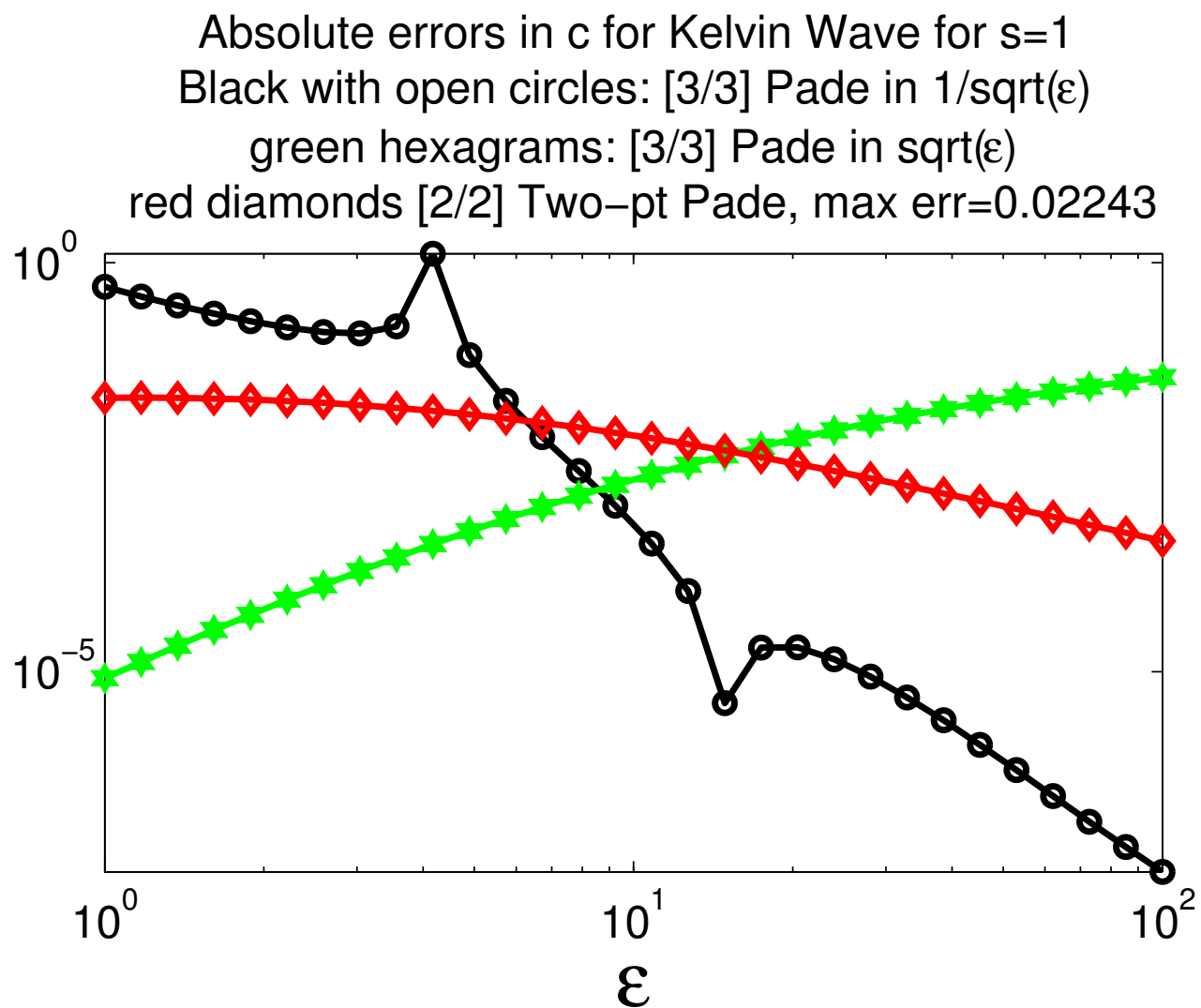


Figure 1: Errors in the small- ϵ and large- ϵ Padé $[3/3]$ approximations and also the quadratic-over-quadratic $c_{[2/2]}^{two-point}$ two-point rational approximation that for the phase speed for the Kelvin mode for $s = 1$.

Deriving Asymptotic Series by Galerkin Methods & Computer Algebra

- Galerkin Matrix Elements by Exact, Analytical Integration
 - Hermite function basis [large ϵ]
 - spherical harmonic basis [Small ϵ]
- Expand in ϵ or $1/\sqrt{\epsilon}$ & match powers
- Solve order-by-order in exact rational arithmetic

Low order small ϵ expansions by Dikii & Golitsyn and by Longuet-Higgins circa 1965
LH gave limited results for large ϵ

Exponential Smallness & Hermite Functions

[Define $\mu = \sin(\textit{latitude})$]

Key step in large ϵ , Hermite function asymptotics is

$$\boxed{\mu \rightarrow \mathcal{Y} / \sqrt{\sqrt{\epsilon}}}$$

Paradox: $\mathcal{Y} \in [-\epsilon^{-1/4}, \epsilon^{-1/4}]$

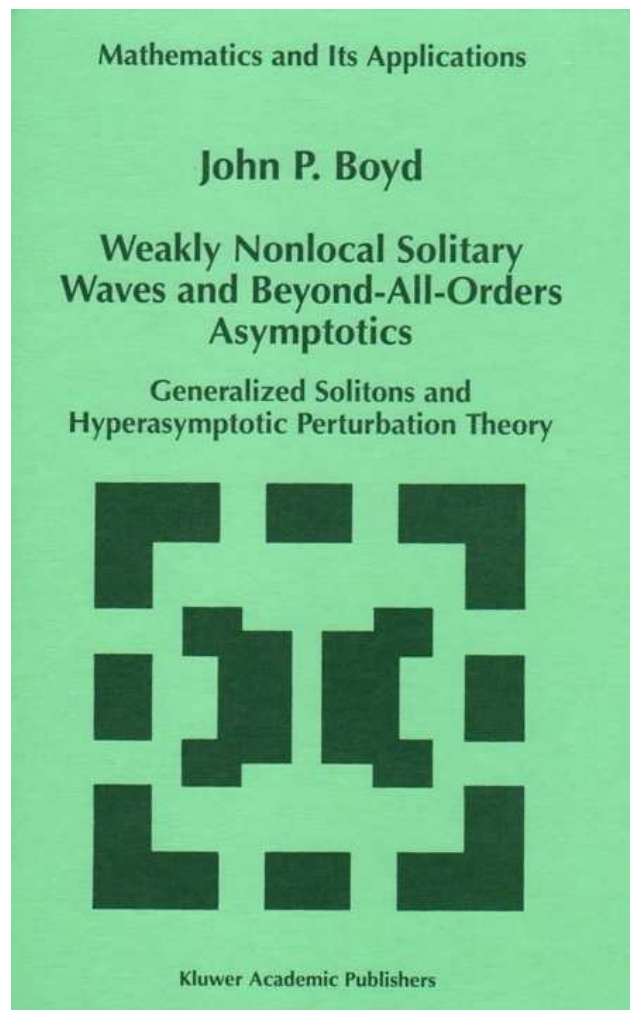
but Galerkin integrals are on $\mathcal{Y} \in [-\infty, \infty]$

- error is EXPONENTIALLY SMALL
- $\exp(-1/\sqrt{\epsilon})$ is INVISIBLE to ϵ -power series
- Coefficients of asymptotic series are EXACT & RATIONAL
- Series DIVERGE

Reviews on Exponential Smallness

"The Devil's Invention: Asymptotics, Superasymptotics and Hyperasymptotics", *Acta Applicandae*, **56**, 1-98 (1999).

"Hyperasymptotics and the Linear Boundary Layer Problem: Why Asymptotic Series Diverge, *SIAM Rev.* , **47**, no. 3, 553-575 (2005)



SUMMARY

- Equatorial trapping depends on $s^2 + \epsilon$
[zonal wavenumber (squared) plus
Lamb's parameter]
- Kelvin Cnoidal Wave/Corner Wave/Breaking:

Small amplitude Kelvin:

cnoidal waves & solitons

Largest non-breaking Kelvin wave is
a corner wave

Medium & large amplitude Kelvin:
frontogenesis and breaking

- Hough point clouds can be connected by per-
turbation series and two-point Padé approx-
imations

In preparation: "Hough Functions: Revisit-
ing Longuet-Higgins' Masterwork Half a Cen-
tury Later"